



Credit: CTA

# Reviving Stellar Intensity Interferometry with the Cherenkov Telescope Array: Laboratory simulation of stellar observations

---

**Tiphaine Lagadec - Lund Observatory, Sweden**  
lagadec.tiphaine@gmail.com

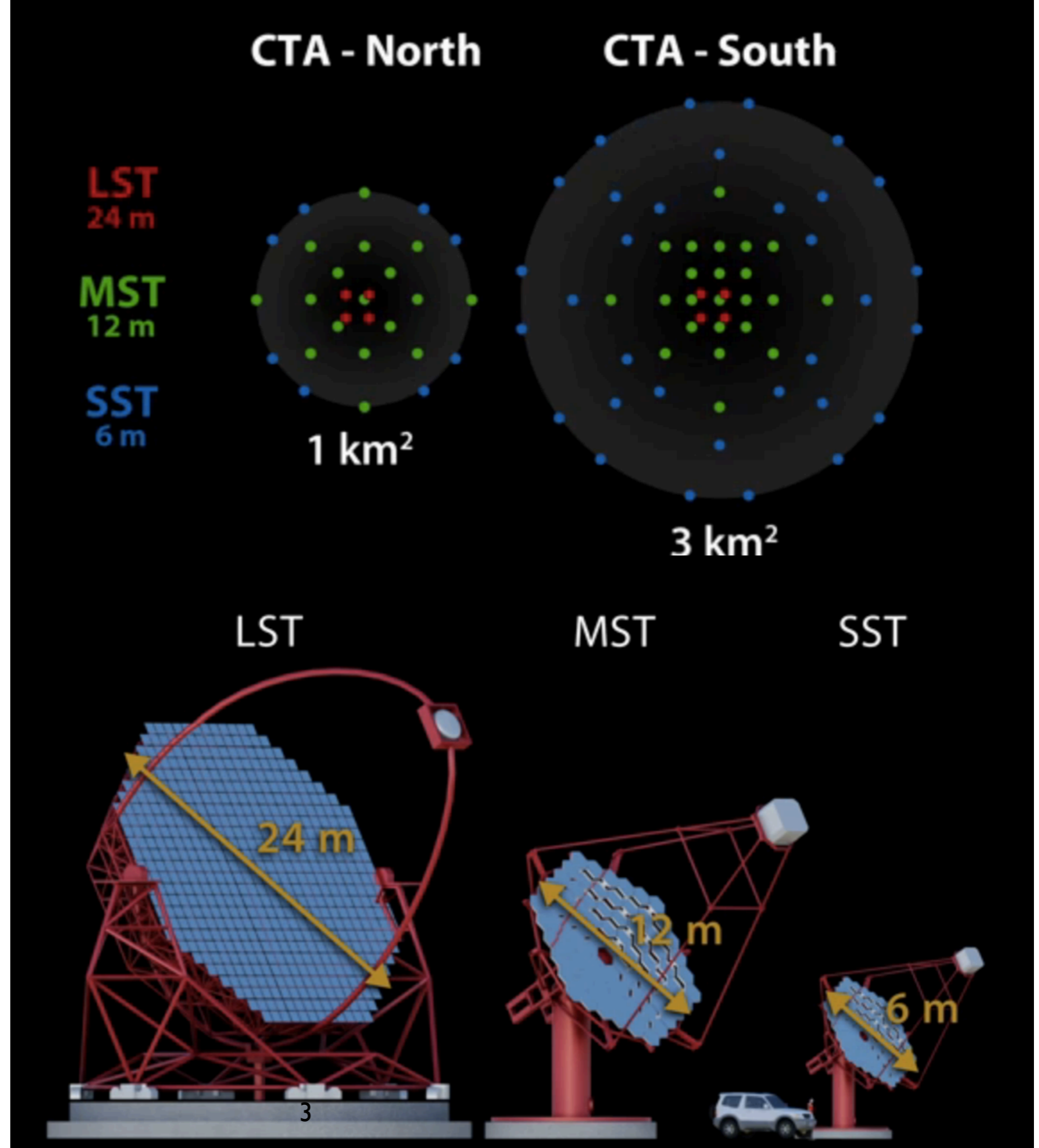
# Content

- *A long held dream...*
- *Experimental simulations*
- *Method*
  - How to make an artificial star in the lab?*
  - The detectors*
  - The correlation*
- *Results*
  - 1D angular measurements*
  - 2D angular maps*
  - 2D binary star*
- *Pending issues*

2 arrays  
→ Full sky coverage

km baseline  
→  $\sim \mu\text{as}$  resolution

many baselines  
→ covers uv-plane



## *Experimental simulation*

# **End-to-end experimental simulation of SII**

To be simulated:

- Multiple baselines
- Complex stellar structures

To be tested:

- Digital correlators
- Semiconductor detectors

Overall aim:

*Demonstrate the procedure of SII with CTA*

- Retrieve correlations between pairs of telescopes
- Retrieve signatures of complex stellar structures
- Retrieve images



*The method*  
*How to make a star in the lab?*

Requirements:

- Small angular diameter
- Complex structures
- Bright and hot
- Chaotic thermal light (Gaussian amplitude distribution)

*The method*  
*How to make a star in the lab?*

Requirements:

- Small angular diameter

Illuminated pinholes

100  $\mu\text{m}$  pinhole @ 23 m  $\Rightarrow$  1 arcsecond star

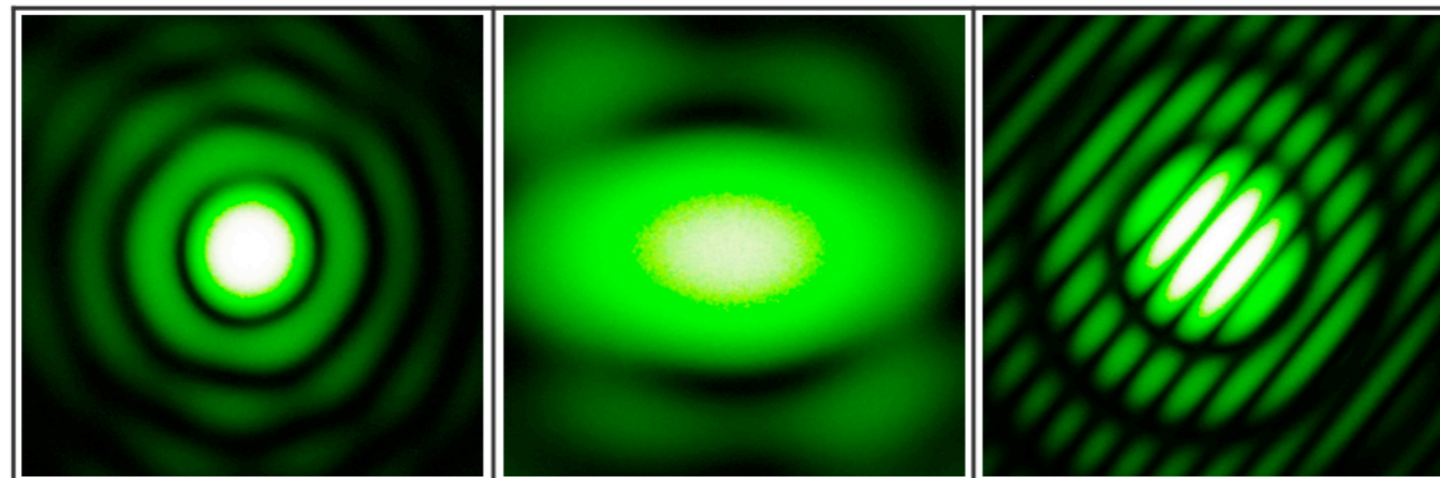
*The method*  
*How to make a star in the lab?*

Requirements:

- ~~Small angular diameter~~
- Complex structures

Pinholes come in different shapes and sizes  
(round, elliptical, double aperture)

⇒ Simulating “realistic” star shapes and angular diameters



*The method*  
*How to make a star in the lab?*

Requirements:

- ~~Small angular diameter~~
- ~~Complex structures~~
- Bright and hot

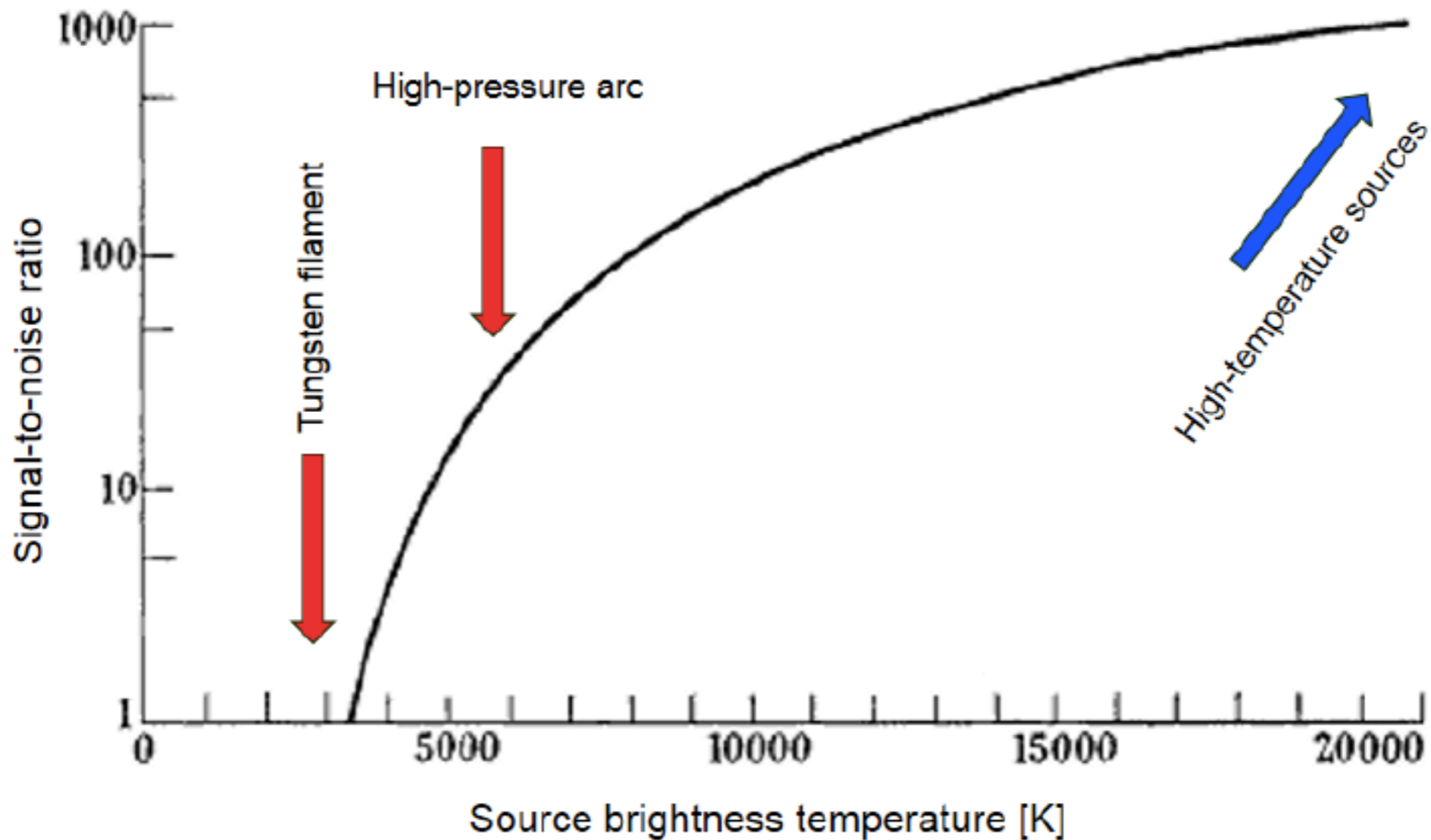


# The method

## How to make a star in the lab?

### The problem of the temperature

SNR as a function of the source brightness temperature from R.Hanbury Brown and R.Q.Twiss:



*The method*  
*How to make a star in the lab?*

Requirements:

- ~~Small angular diameter~~
- ~~Complex structures~~
- ~~Bright and hot~~

Laser radiation very bright

*The method*  
*How to make a star in the lab?*

Requirements:

- ~~Small angular diameter~~
  - ~~Complex structures~~
  - ~~Bright and hot~~
  - Chaotic thermal light (Gaussian amplitude distribution)
- ? Laser radiation very bright **but not thermal** !  
(no intensity fluctuations in space and time)

*The method*  
*How to make a star in the lab?*

Requirements:

- ~~Small angular diameter~~
- ~~Complex structures~~
- ~~Bright and hot~~
- Chaotic thermal light (Gaussian amplitude distribution)

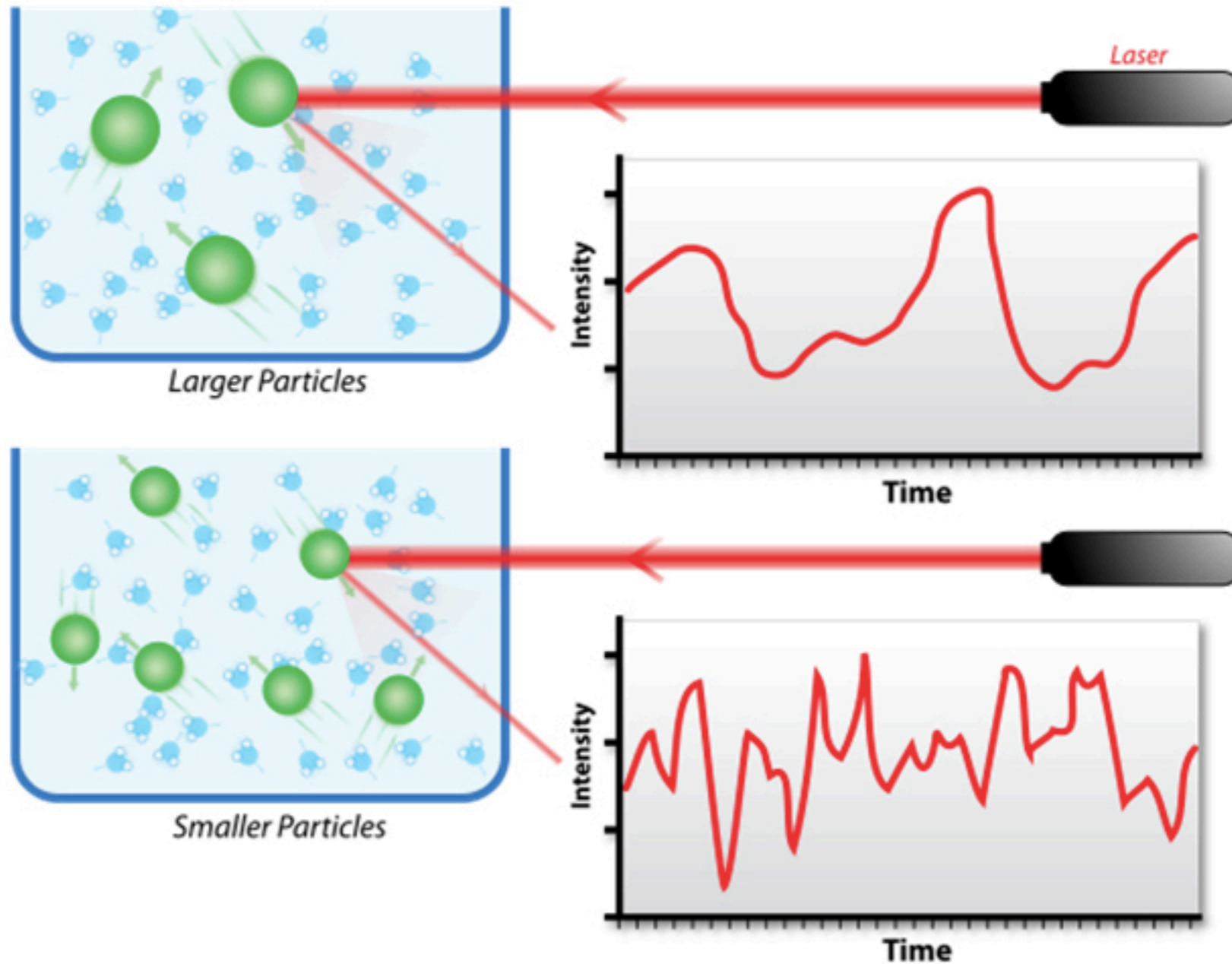
Dynamic Light Scattering (DLS):

Coherent bright radiation **scattered**  
Resulting radiation is thermal and Doppler broadened



*The method*  
*How to make a star in the lab?*

Dynamic Light Scattering



# The method

## How to make a star in the lab?

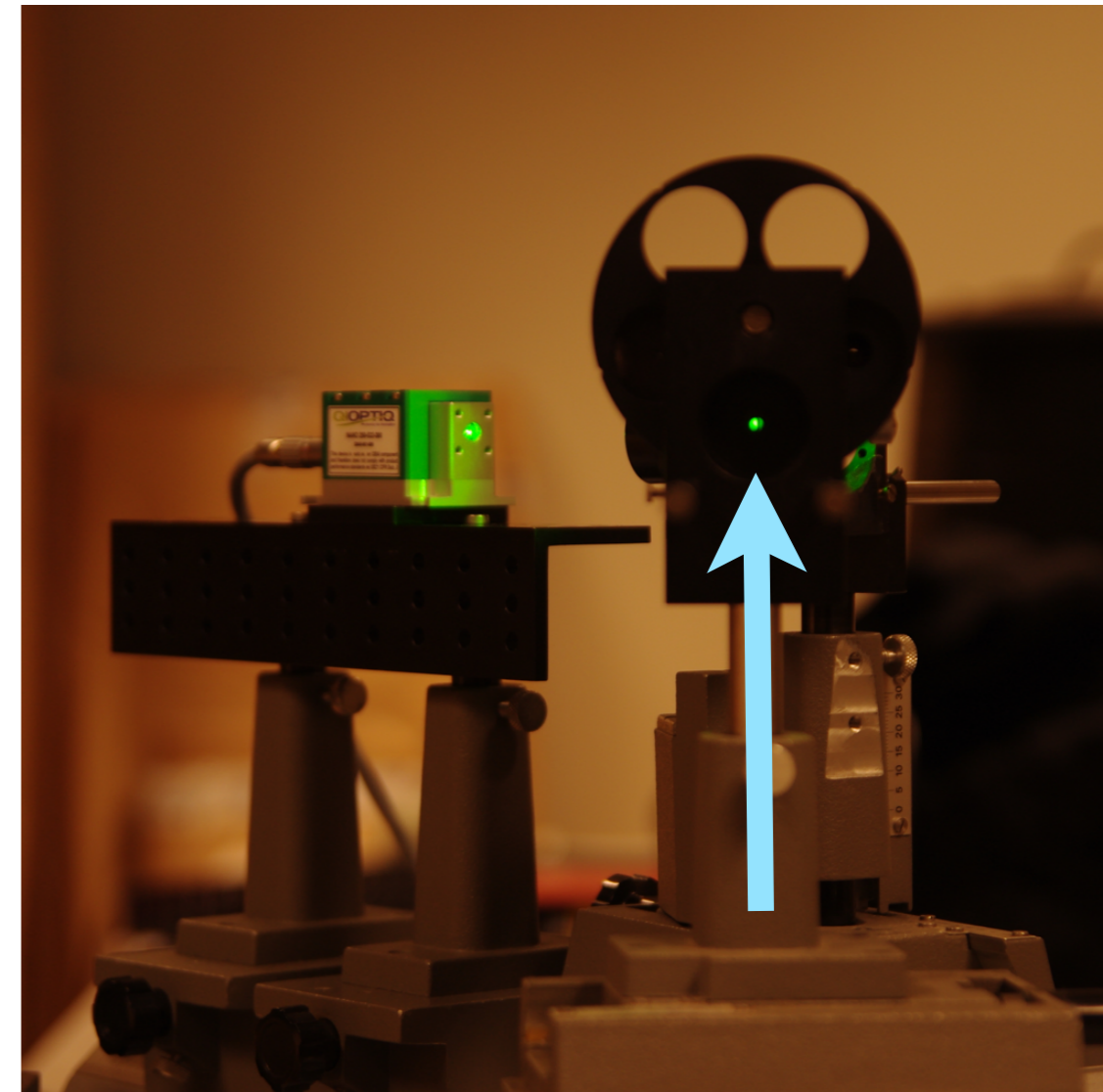
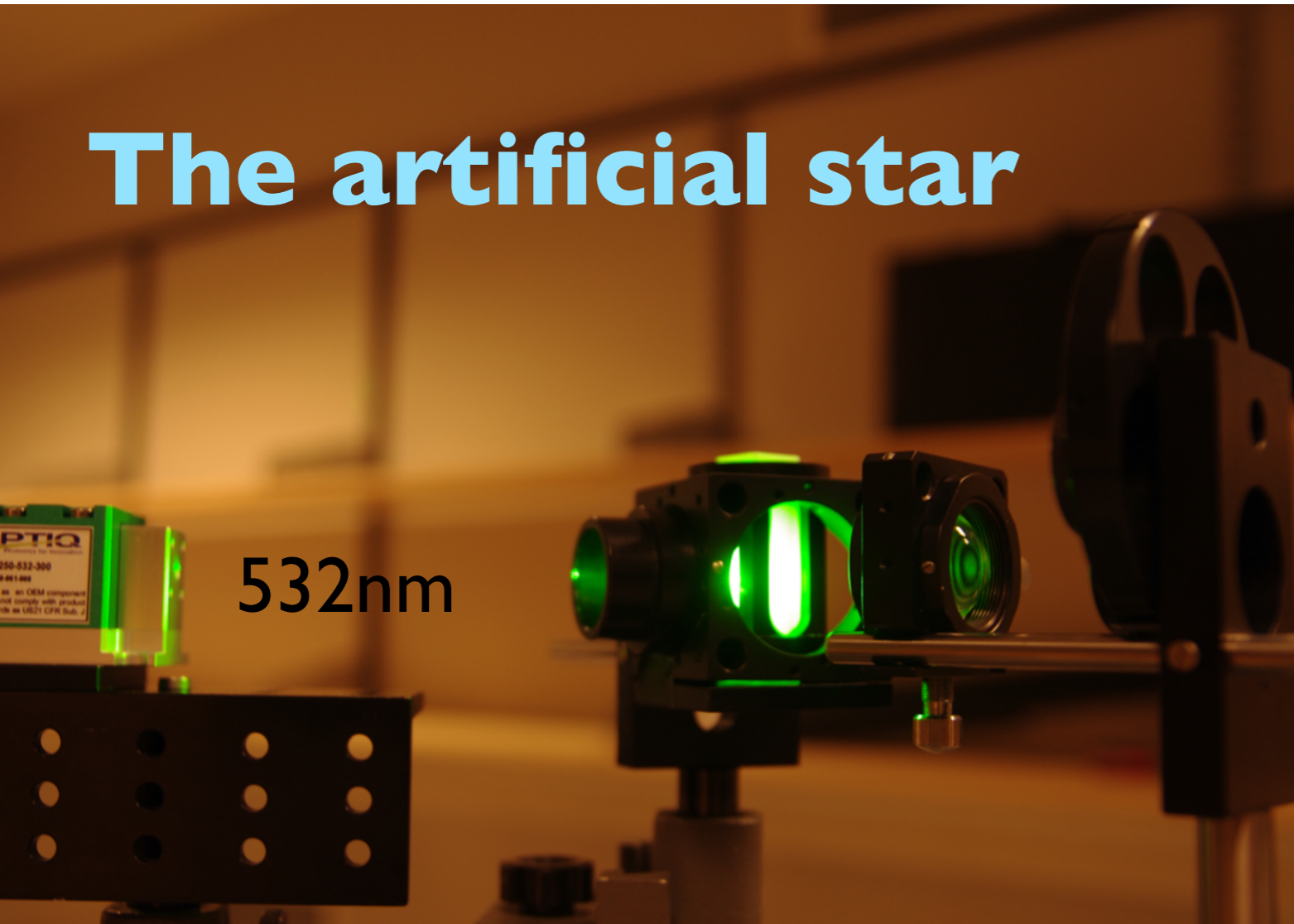
### Requirements:

- ~~Small angular diameter but good photon fluxes~~
- ~~Complex structures~~
- ~~Bright and hot~~
- ~~Chaotic thermal light (Gaussian spectrum)~~

⇒ We have our artificial star !

*The method*  
*How to make a star in the lab?*

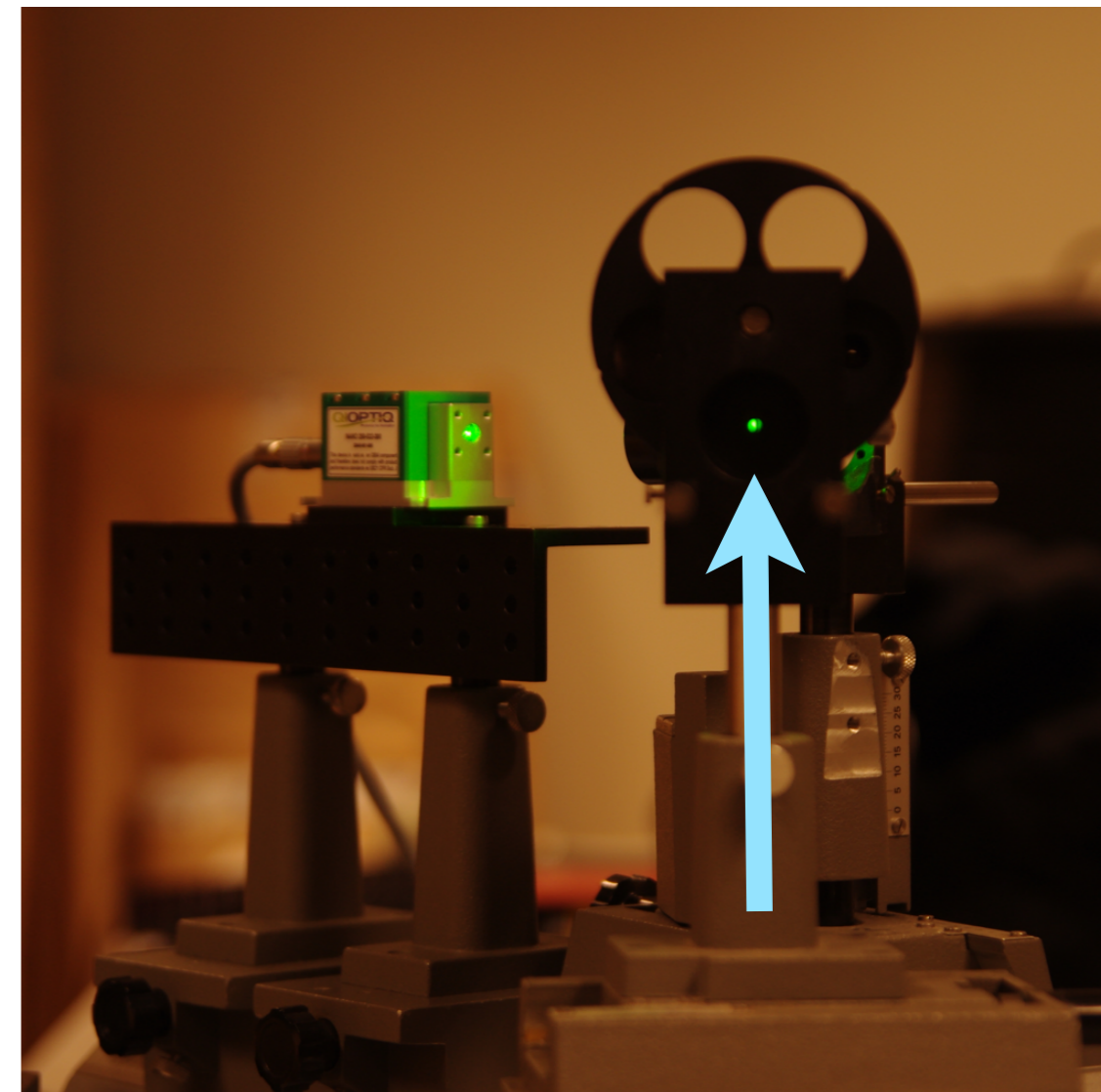
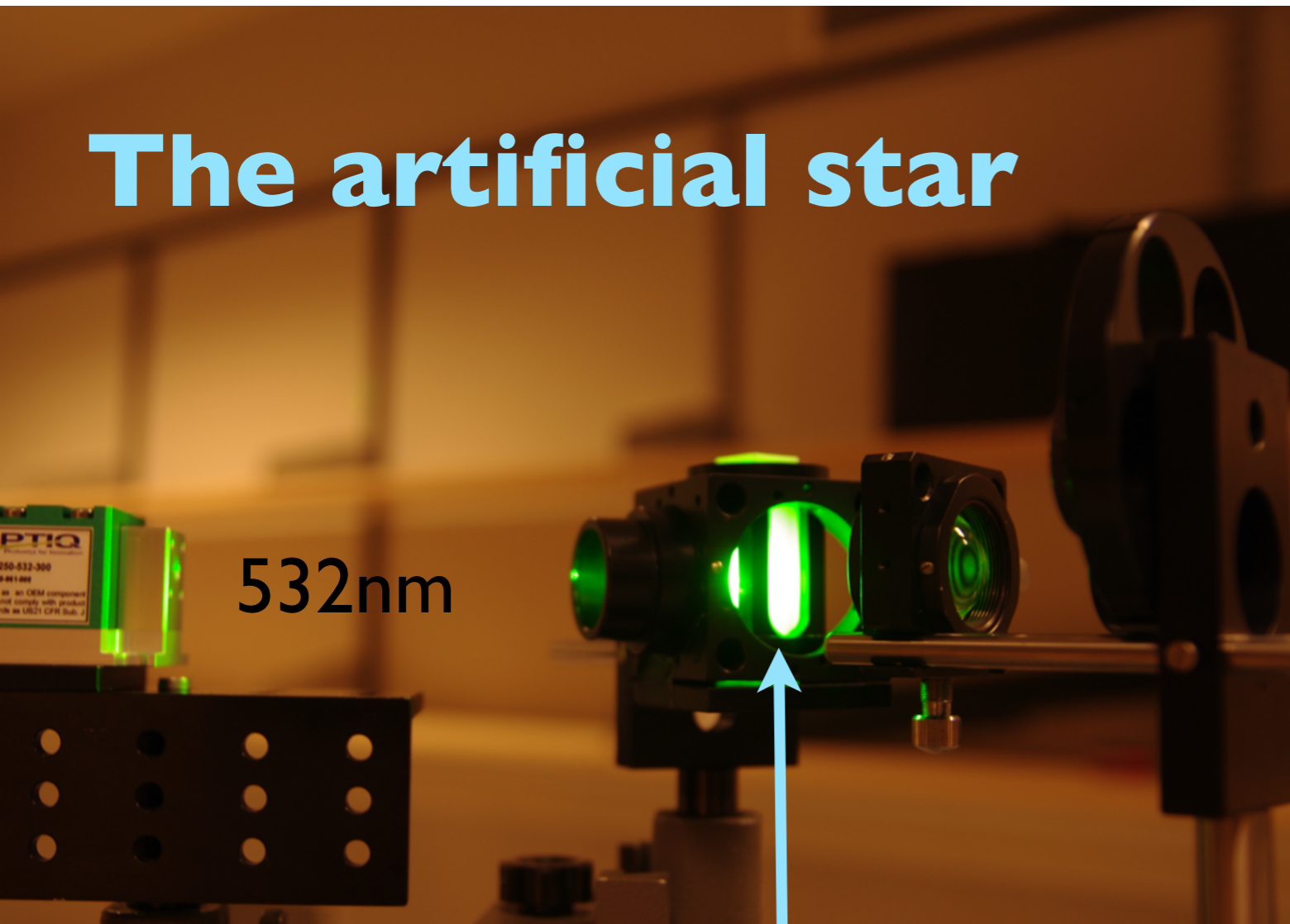
**The artificial star**





*The method*  
*How to make a star in the lab?*

**The artificial star**



Scatterer: plastic micro-spheres or fat globules in milk



*The method*  
*How to make a star in the lab?*

**$\approx 23\text{m}$  away**



*The method*  
*The detectors*

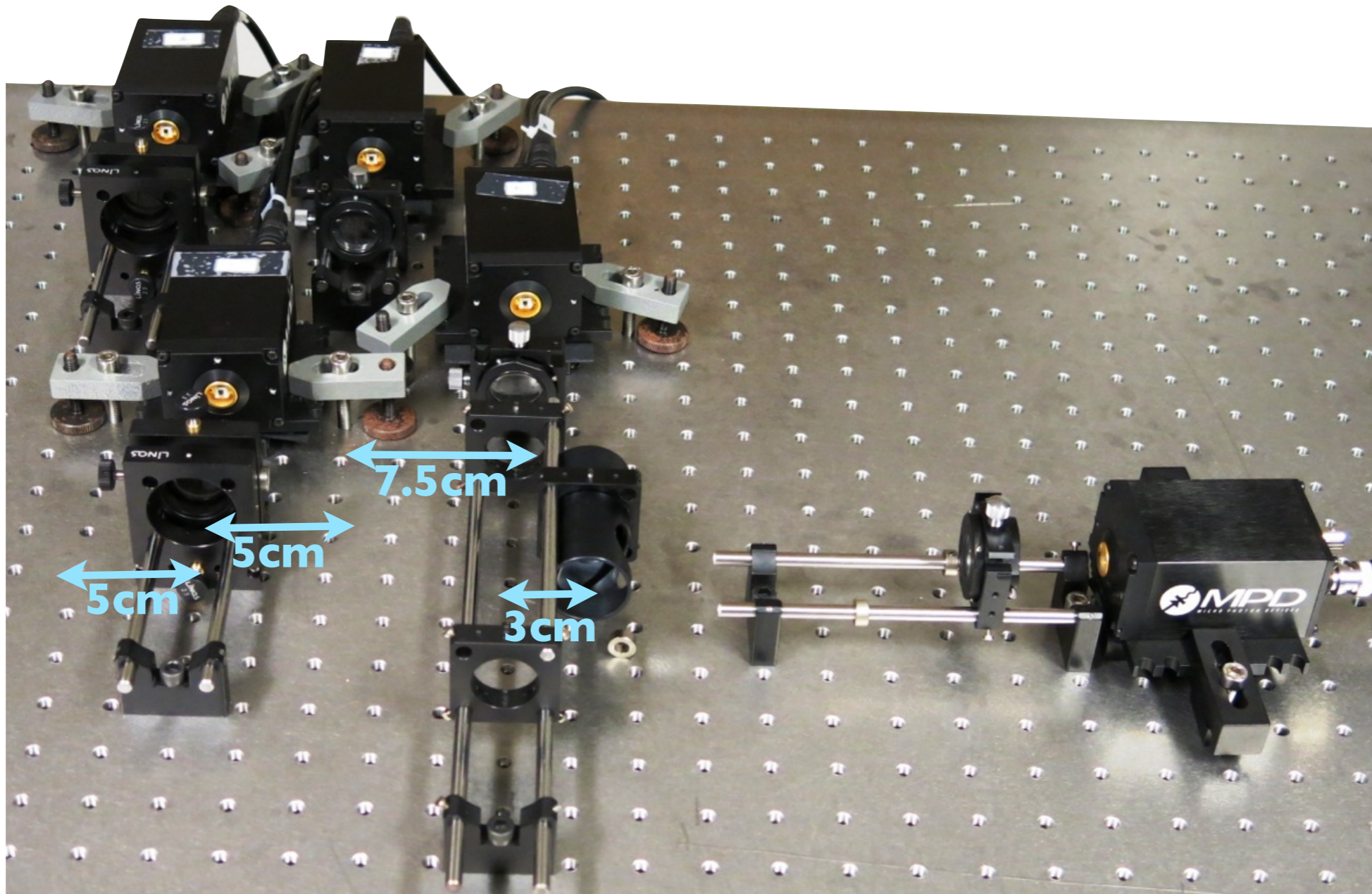
## Single Photon Avalanche Diode (Geiger mode APD)

- Analogous to planned CTA detectors
- Very high QE
- High gain
- Large bandwidth
- Compact
- Fast ( $\sim$ ns)
- Immune to magnetic fields



# The method

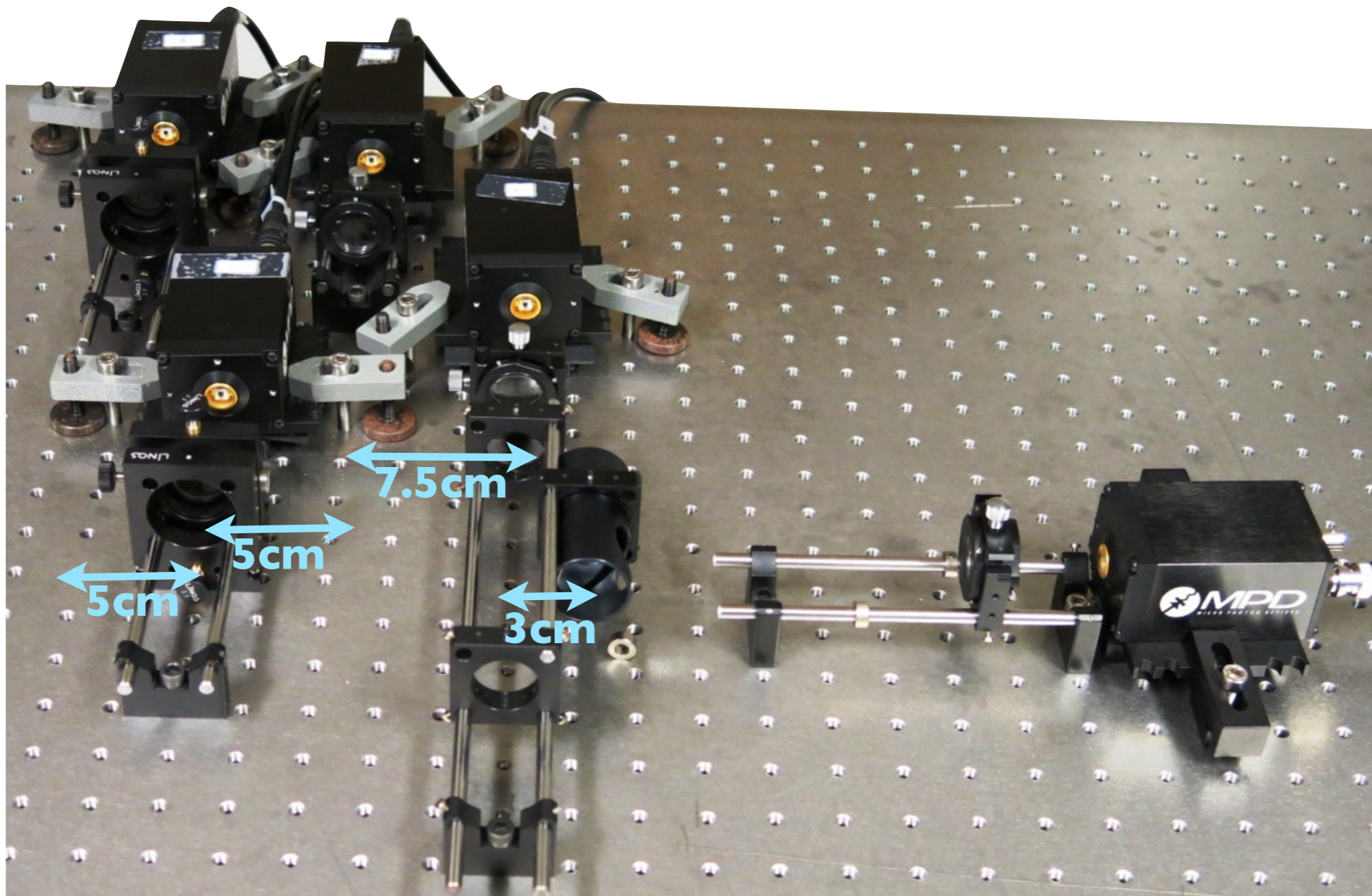
## A 2D telescope array in the lab



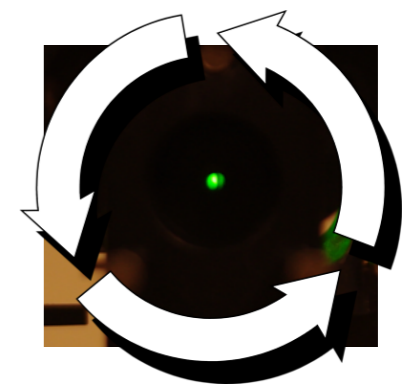


# The method

## A 2D telescope array in the lab

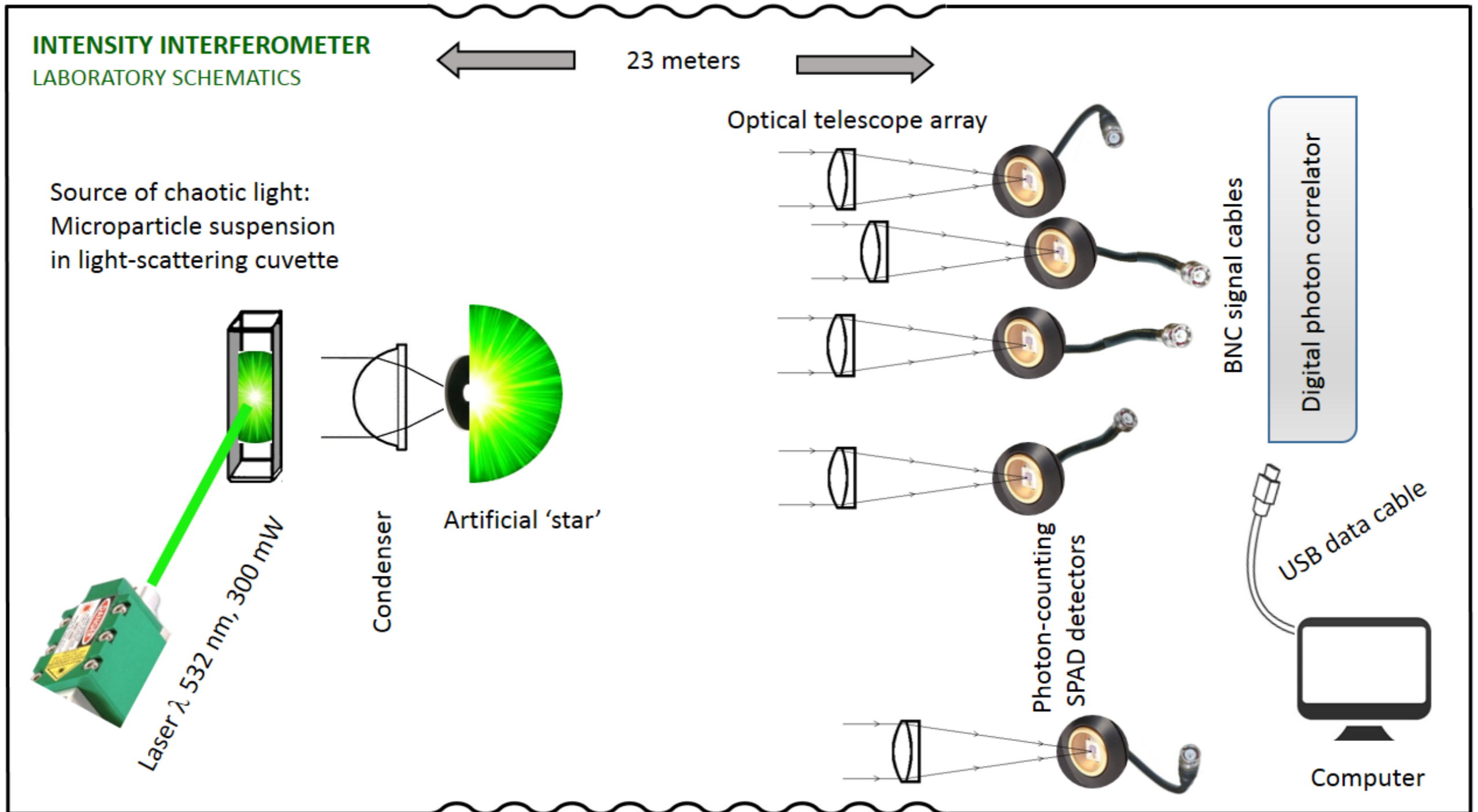


**Rotating source angle enables 2D-array**



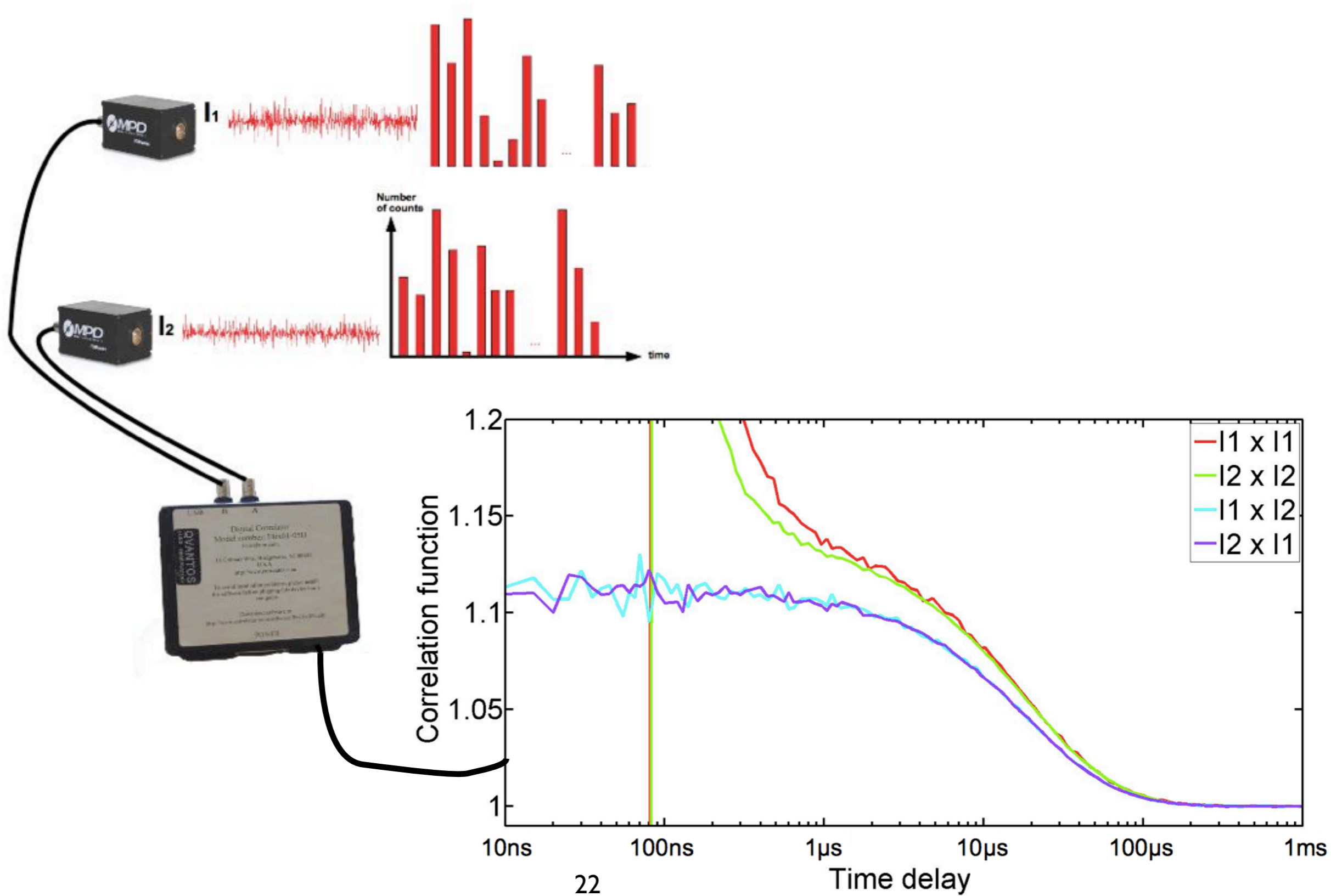
# The method

## The set-up



# The method

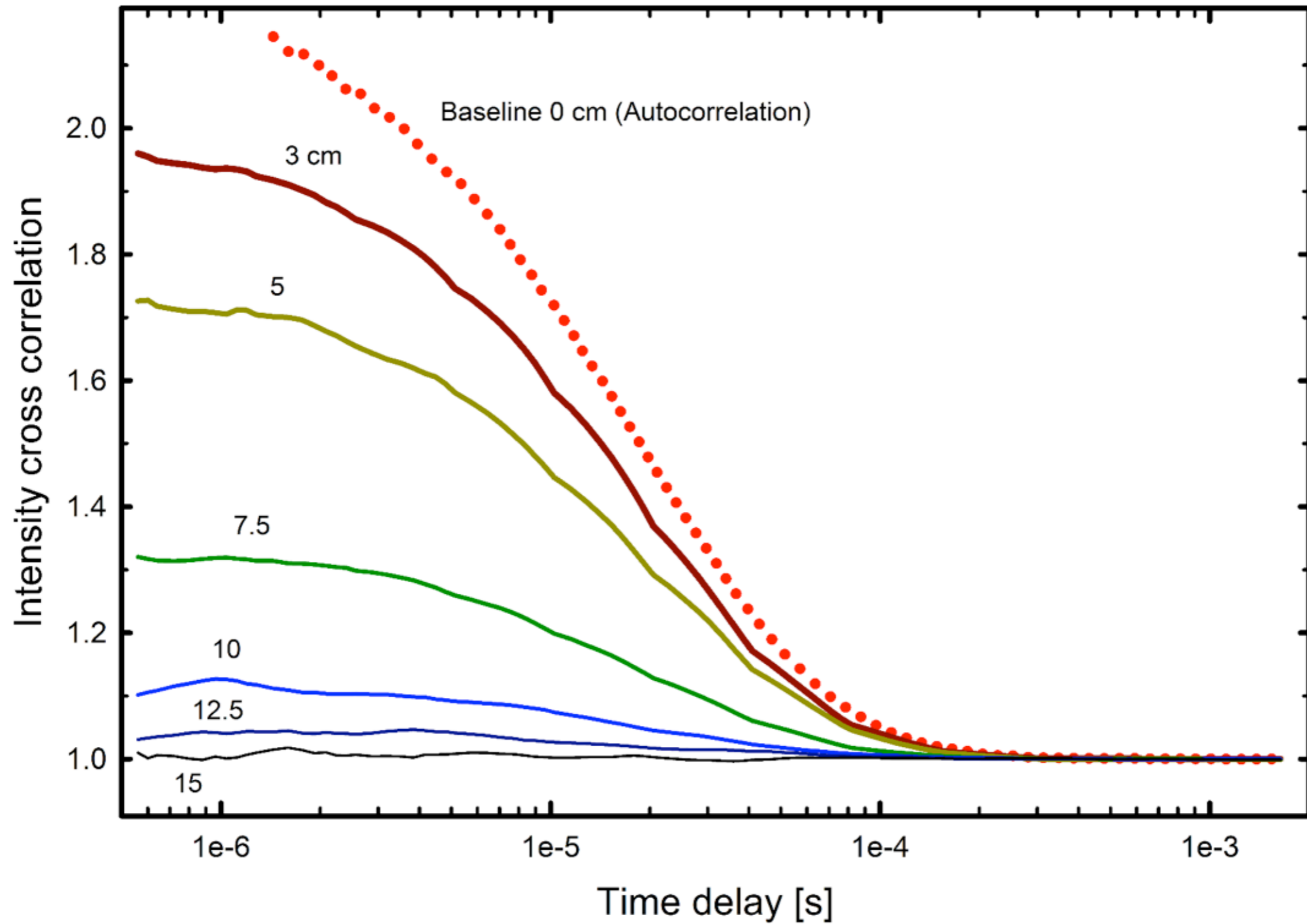
## The correlation



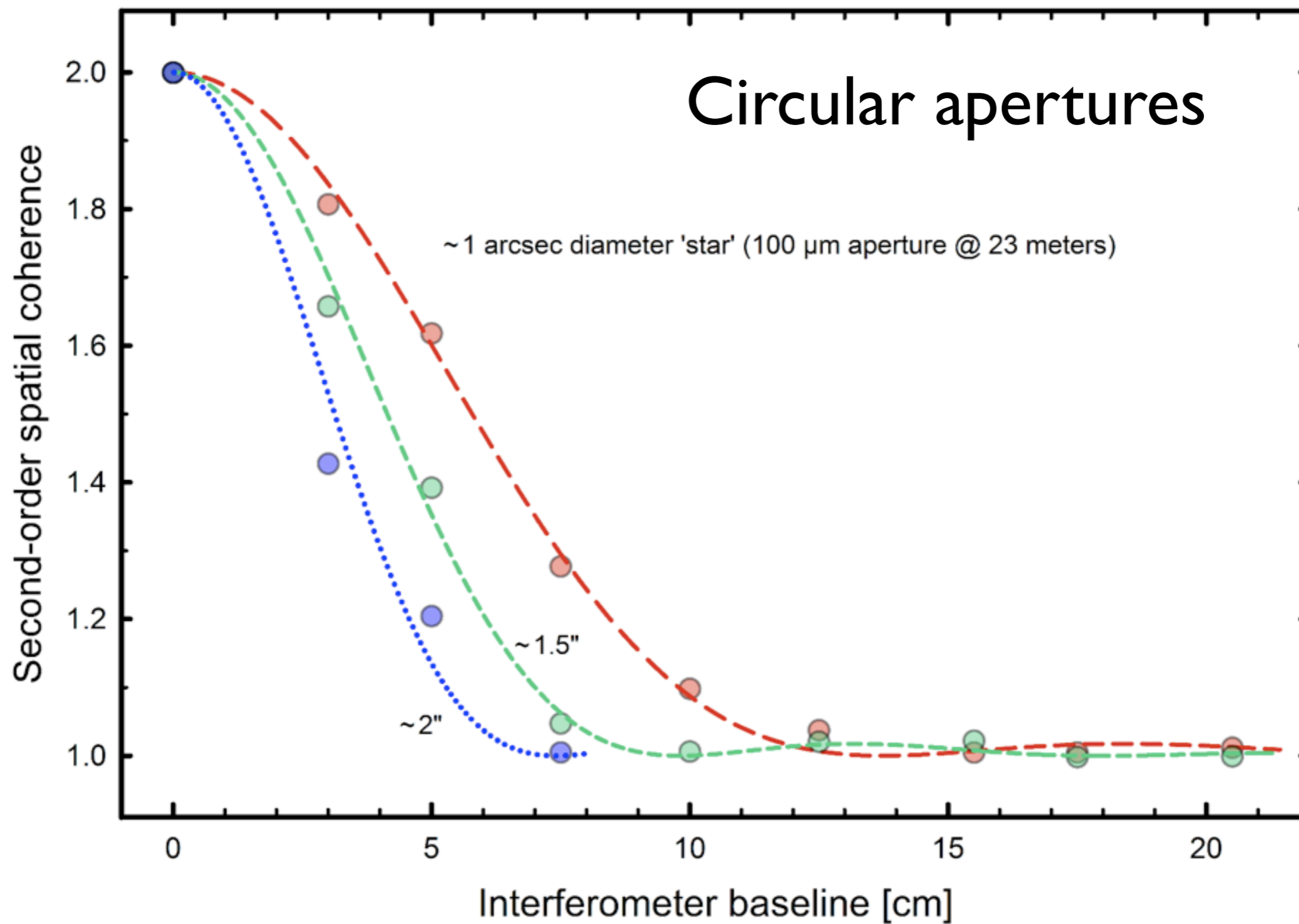


# The results

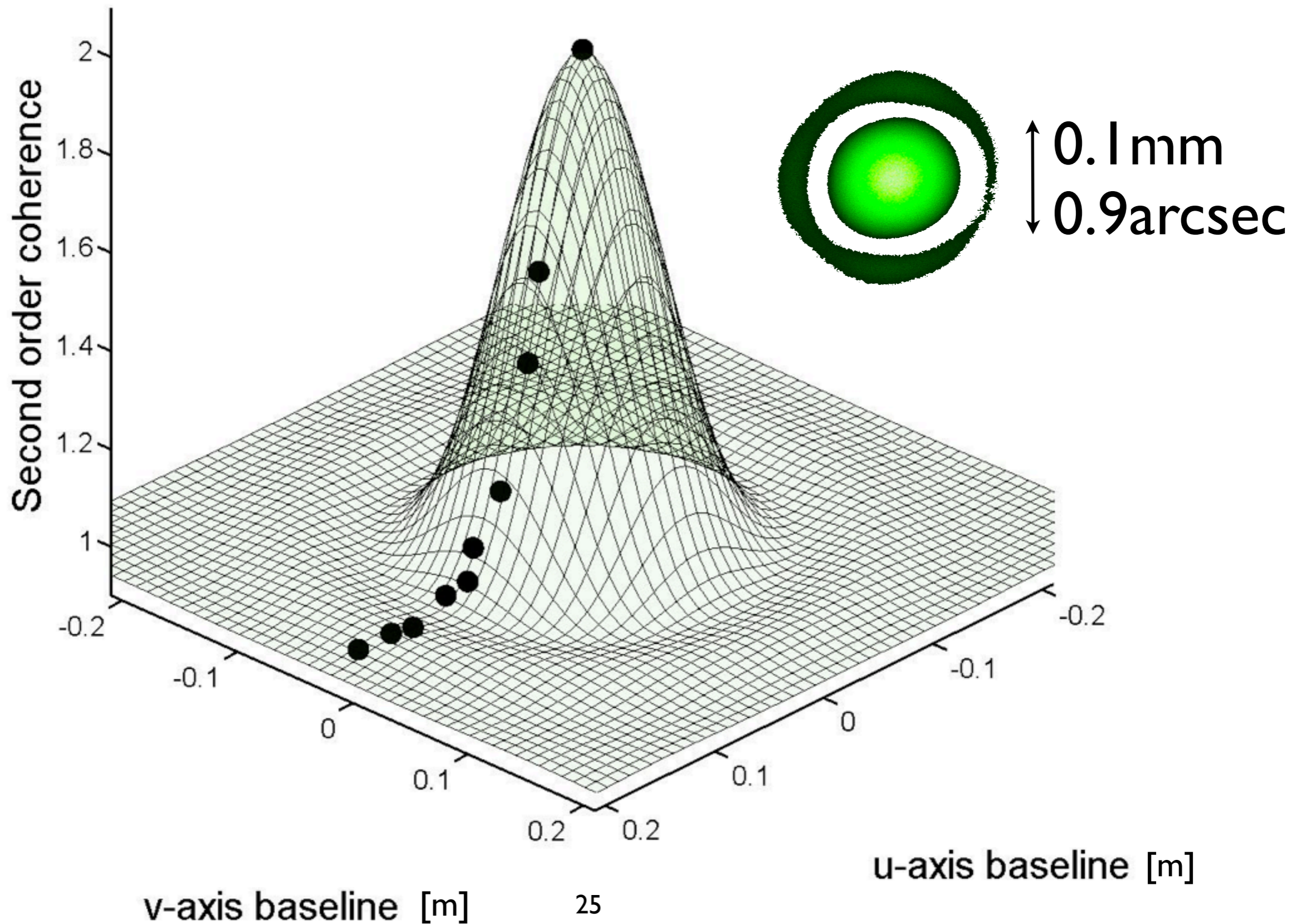
## *ID angular diameter*



*The results*  
*ID angular diameter*



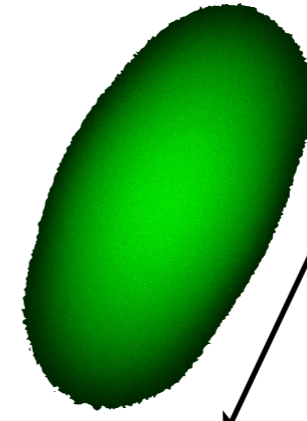
*The results*  
*ID angular diameter*



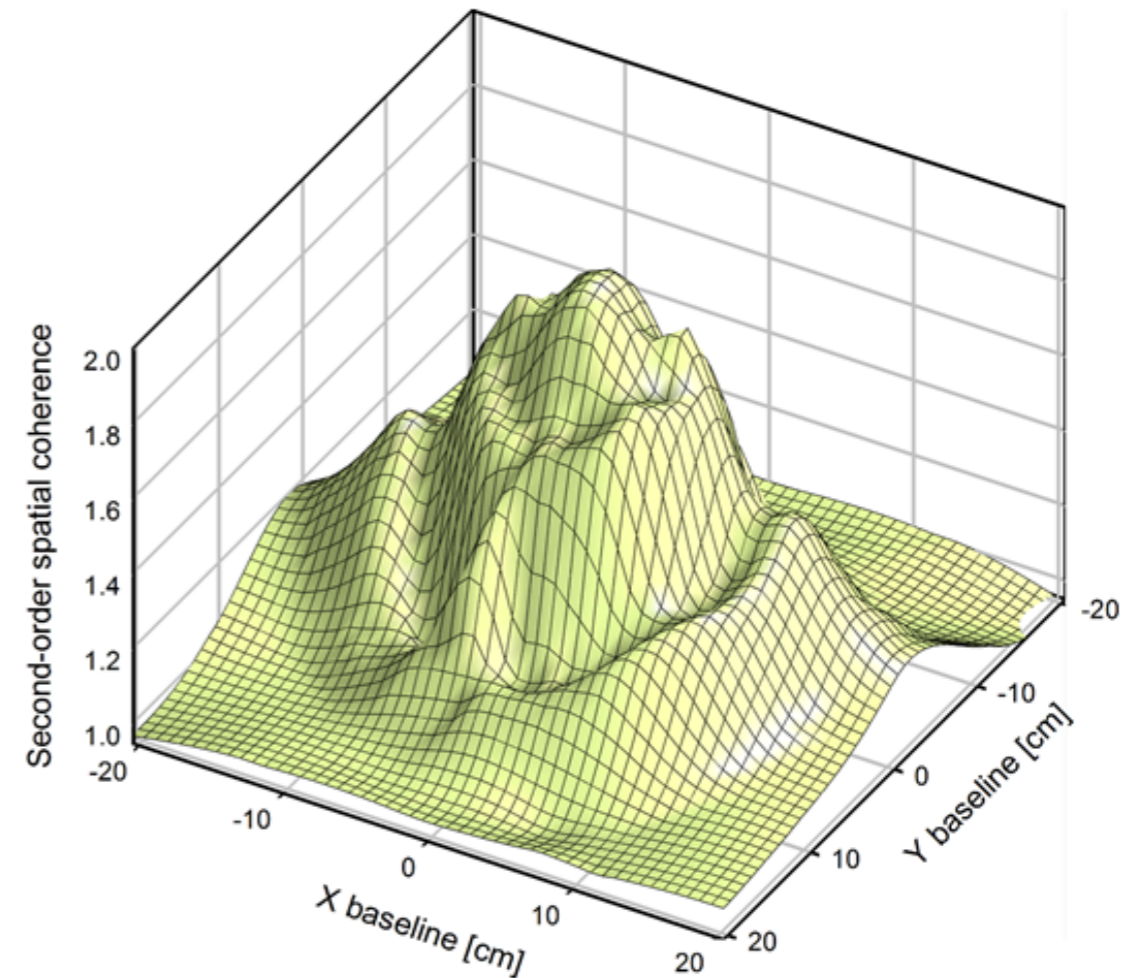
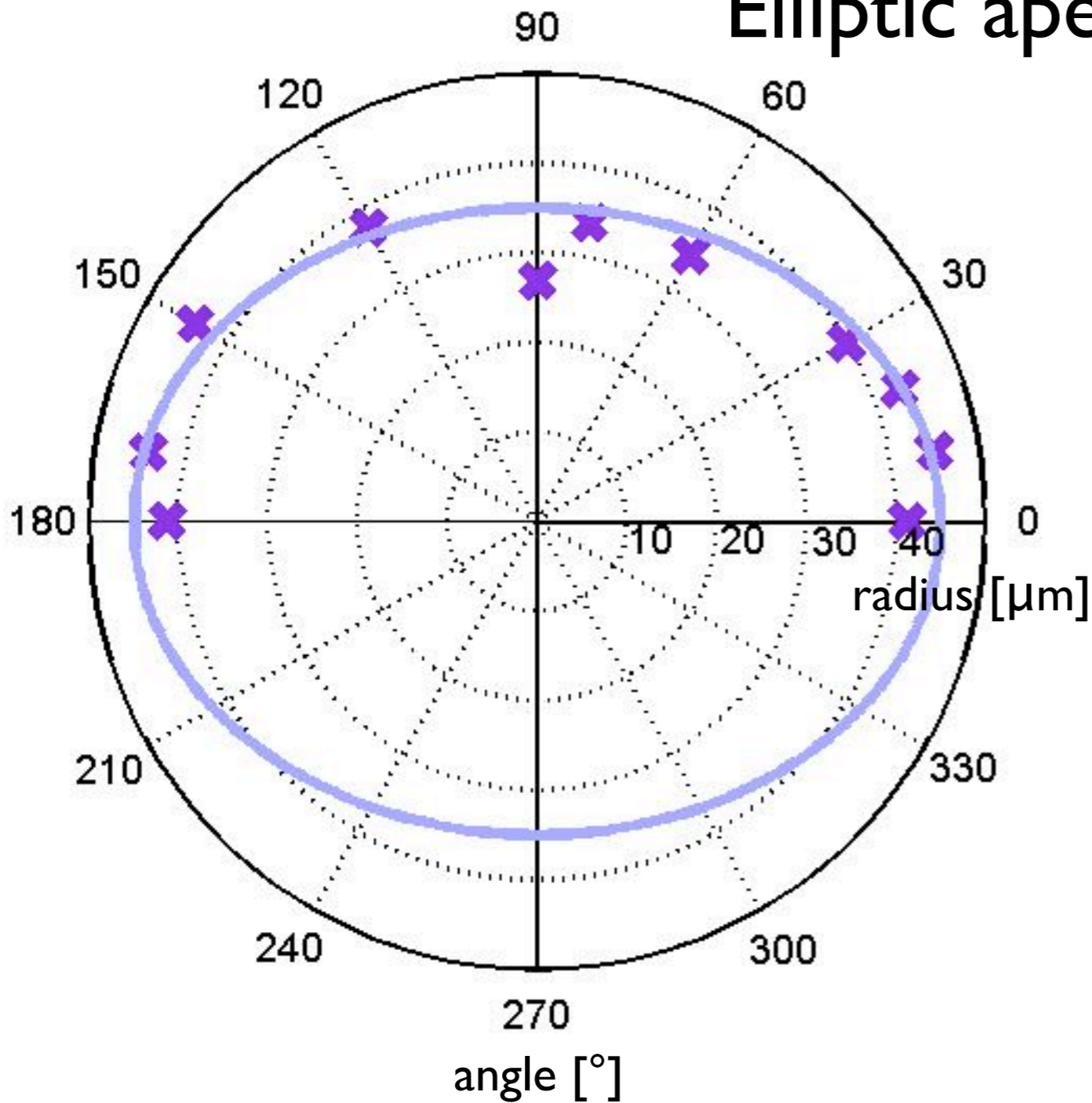


The results  
2D angular diameter

Elliptic aperture



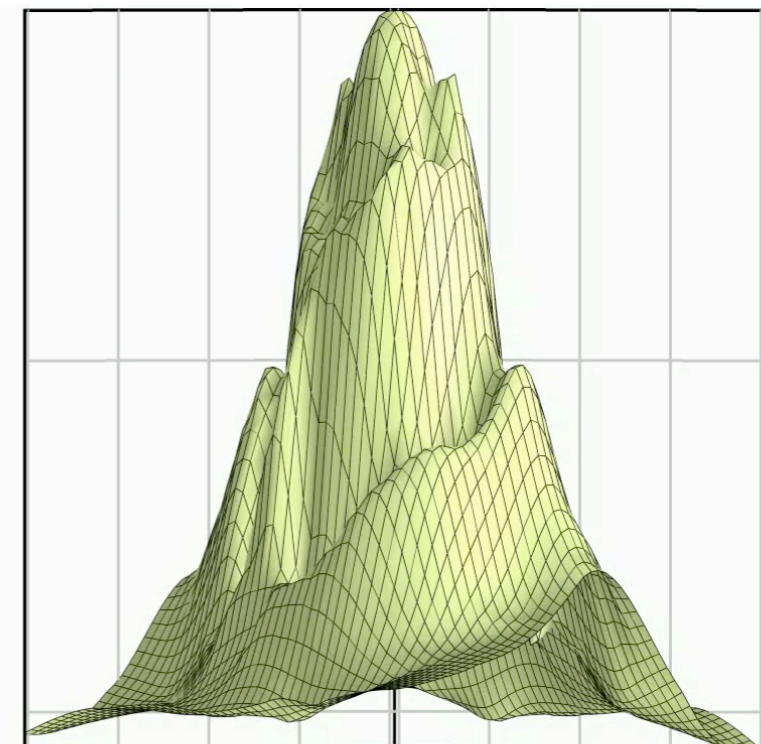
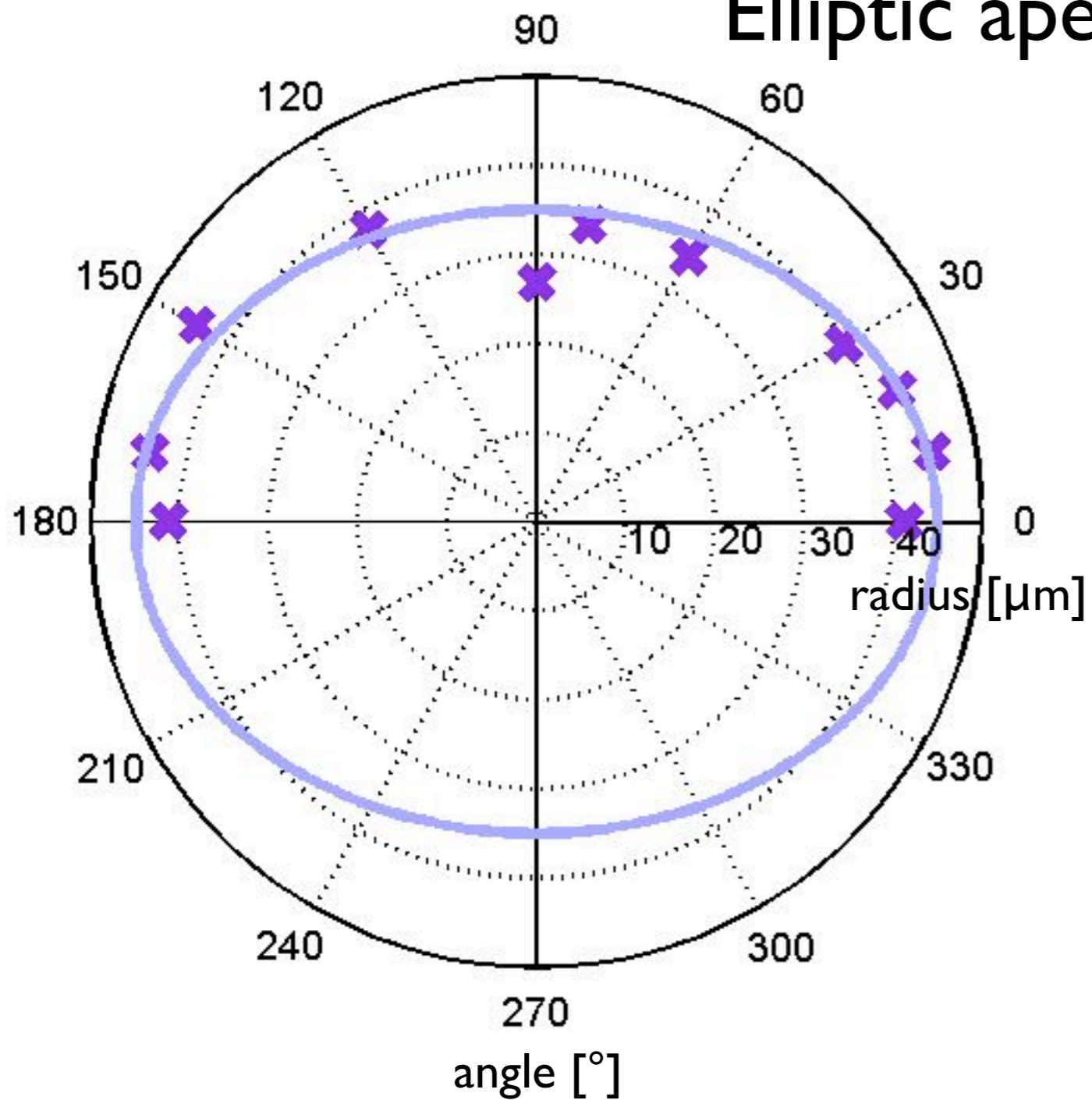
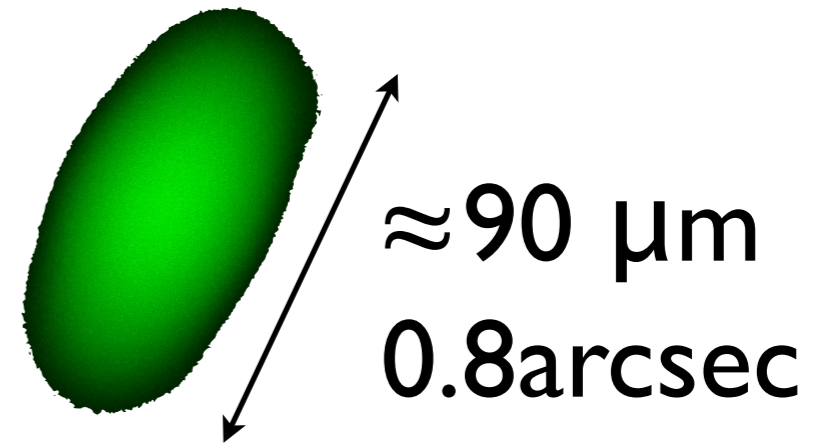
$\approx 90 \mu\text{m}$   
 $0.8 \text{arcsec}$



**100 baselines**

*The results*  
*2D angular diameter*

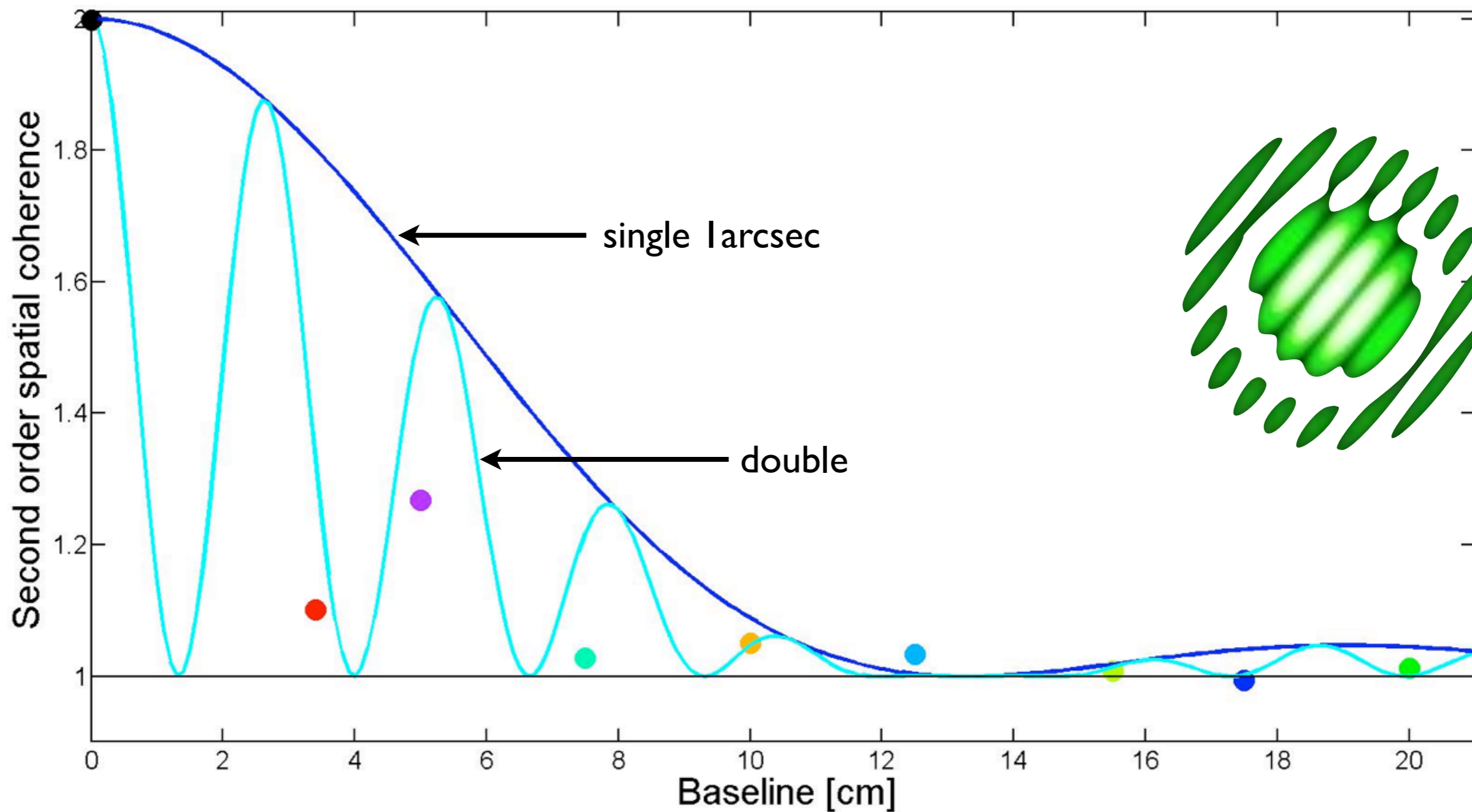
Elliptic aperture



**100 baselines**

*The results*  
*2D binary star*

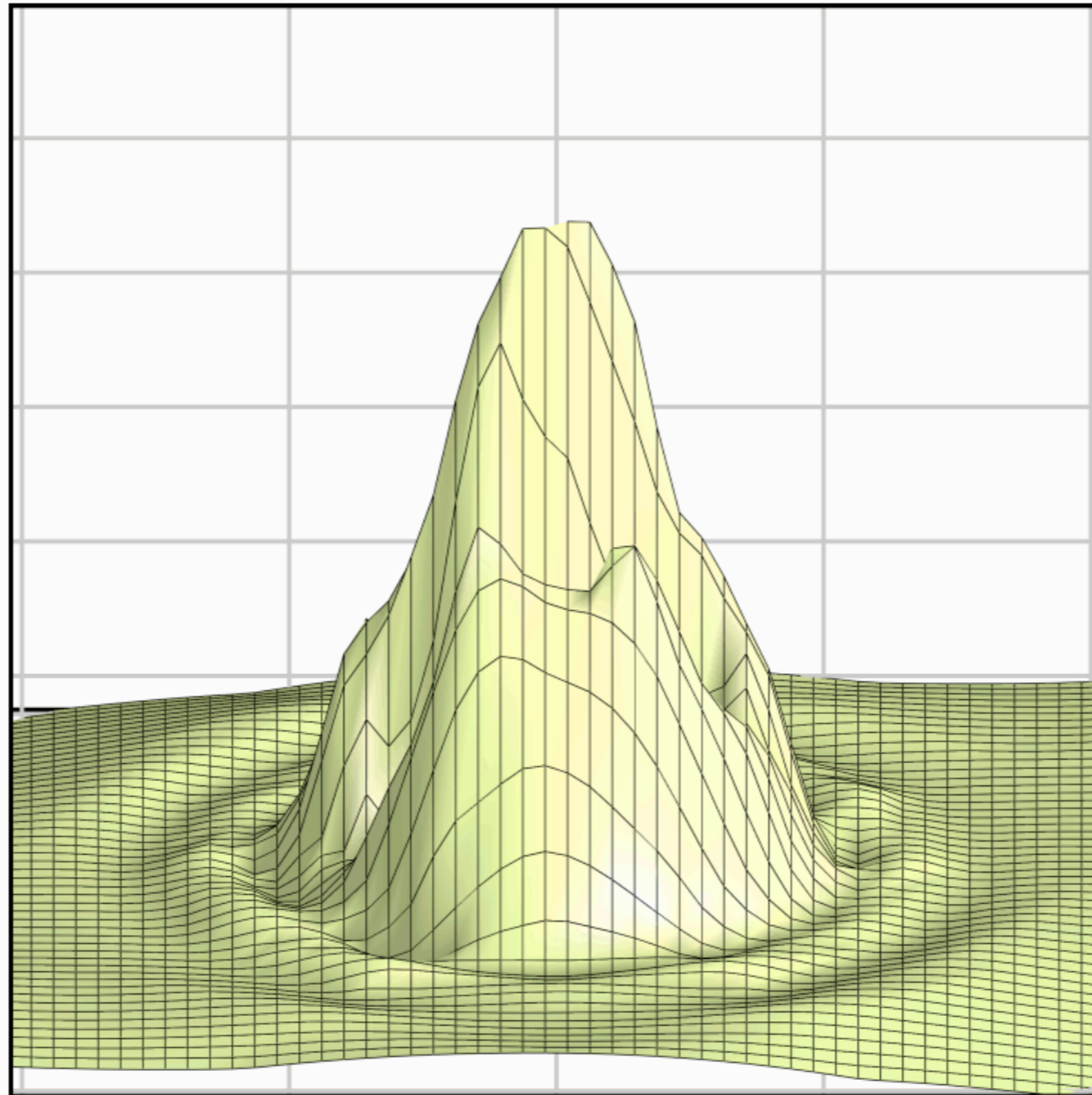
# Double aperture





*The results*  
*2D binary star*

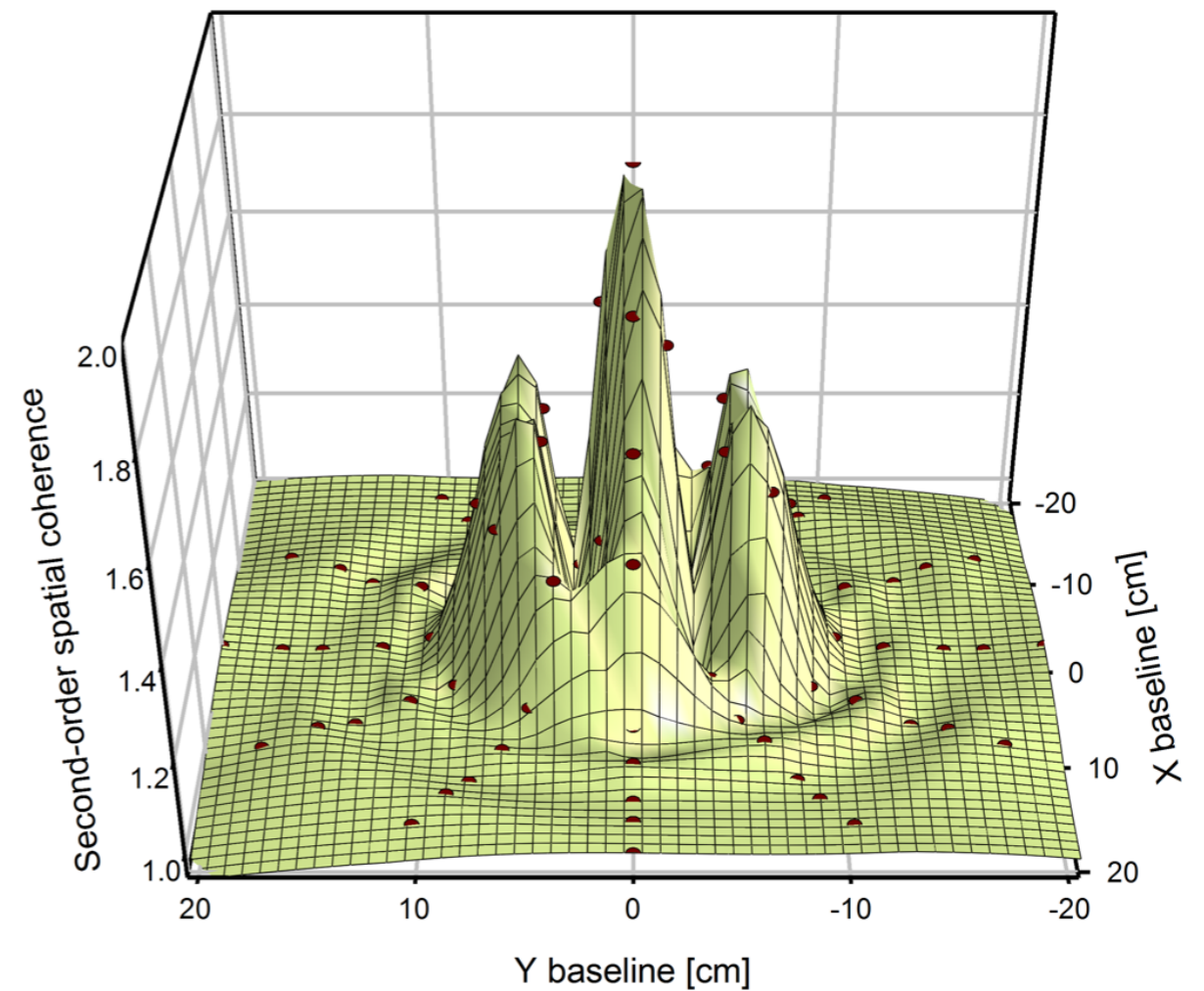
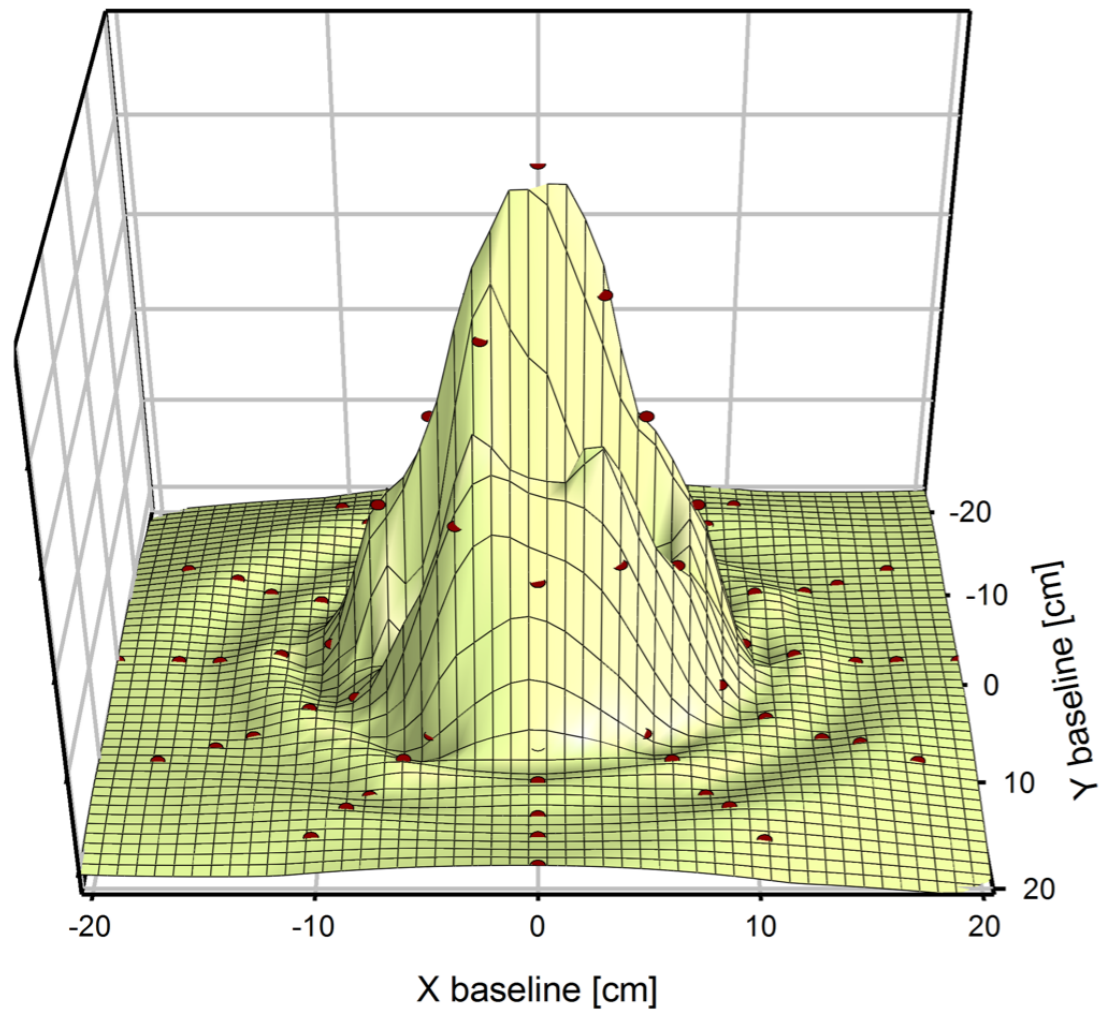
60 baselines



# The results

## 2D binary star

60 baselines





## Pending issues

- *Image restoration*
- *Possible white light illumination ?*
- *Polarization effects ?*
- *Multiple spectral pass-bands ?*

**For details see:**

Dainis Dravins & Tiphaine Lagadec: “*Stellar intensity interferometry over kilometer baselines: Laboratory simulation of observations with the Cherenkov Telescope Array*”, in J.K.Rajagopal, M.J.Creech-Eakman & F.Malbet (eds.): “Optical and Infrared Interferometry IV”, SPIE Proc. 9146 (2014)

**For details see:**

Dainis Dravins & Tiphaine Lagadec: “*Stellar intensity interferometry over kilometer baselines: Laboratory simulation of observations with the Cherenkov Telescope Array*”, in J.K.Rajagopal, M.J.Creech-Eakman & F.Malbet (eds.): “Optical and Infrared Interferometry IV”, SPIE Proc. 9146 (2014)

**Thank you for your attention!**